

Categorizing Models of Cosmic Acceleration

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arXiv:1112.0303 with Eanna Flanagan

Work in Progress with Scott Watson, Minjoon Park, Eva-Maria Mueller, Rachel Bean and Eanna Flanagan

Variety is the Spice of Life

- Quintessence
 - k-essence
 - Brans-Dicke theories
 - Ghost Condensate
 - Extra-dimensions a la UED/RS/ADD/DGP
 - $f(R)$ gravity
 - Gauss-Bonnet gravity
 - ...
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- In a low energy four-dimensional limit, all these theories essentially behave as GR + scalar field(s)
 - Perhaps we can perform a general analysis in this regime?

Systematic Characterization

- Effective Field Theory of a scalar field (eg, Park *et al.*, JB and Eanna Flanagan)
- Effective Field Theory of scalar field perturbations (eg, Creminelli *et al.*)
- Effective Field Theory of general modifications to gravity (eg, Battye and Pearson)
- Parameterized modifications to gravity (eg, Baker *et al.*)
- Investigations of completely generic scalar field models (eg, Charmousis *et al.*)

Our Approach

Leading Order Action: GR + Canonical Scalar (Quintessence) Field

$$S_0 = \int d^4x \sqrt{-g} \left\{ \frac{m_p^2}{2} R - \frac{1}{2} (\nabla\phi)^2 - U(\phi) \right\} + S_{\text{matter}} \left[e^{\alpha(\phi)} g_{\mu\nu}, \{\psi\} \right]$$

Perturb the Action

$$\phi, g^{\mu\nu}, R_{\mu\nu\sigma\lambda}, \epsilon_{\mu\nu\sigma\lambda}, T_{\mu\nu}, \nabla_\mu, \square \dots$$

Rules of Analysis

- Expand using a derivative expansion
- Require regime of validity such that perturbative terms are small across cosmological history
- Expansion parameter given by

$$\frac{H_0^2}{M^2} \ll 1$$

for some mass scale M

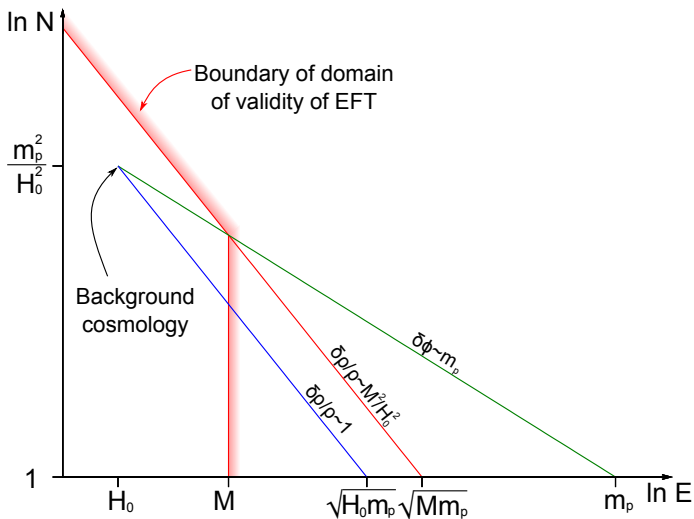
- Remove higher order derivatives in equations of motion (“reduce” the action)
- Impose the Weak Equivalence Principle (Note: not a symmetry of the theory)

Results

$$\begin{aligned}
 S = \int d^4x \sqrt{-g} \left\{ \right. & \frac{m_p^2}{2} R - \frac{1}{2} (\nabla\phi)^2 - U(\phi) + a_1 (\nabla\phi)^4 \\
 & + b_2 T (\nabla\phi)^2 + c_1 G^{\mu\nu} \nabla_\mu \phi \nabla_\nu \phi \\
 & + d_3 \left(R^2 - 4R^{\mu\nu} R_{\mu\nu} + R_{\mu\nu\sigma\rho} R^{\mu\nu\sigma\rho} \right) \\
 & + d_4 \epsilon^{\mu\nu\lambda\rho} C_{\mu\nu}{}^{\alpha\beta} C_{\lambda\rho\alpha\beta} \\
 & + e_1 T^{\mu\nu} T_{\mu\nu} + e_2 T^2 \left. \right\} \\
 & + S_m \left[e^{\alpha(\phi)} g_{\mu\nu} \right]
 \end{aligned}$$

- Coefficients are arbitrary $O(1)$ functions of ϕ (with appropriate dimensionful scalings)
- Parameter space is given by nine free functions

Regime of Validity



Effectiveness of the EFT Approach

- Can describe background and perturbative evolution of the cosmology
- Radiative corrections under control, given constraints on UV theory
- Within regime of validity, yields a very general description

Perturbative Analysis

- Our framework useful for considering background evolution
- Perturbative analysis more suited to EFT of Inflation treatment (Creminelli *et al.*)
- Translation “dictionary” in progress
- EFT of Inflation framework needs extending to treat dark energy-matter interactions generally

Idea of EFT of Inflation

- Work in Unitary Gauge (scalar eaten by metric) with broken time diffeomorphisms
- Specify background evolution of FRW cosmology
- Construct action as leading order terms + quadratic perturbations

$$S = \int d^4x \sqrt{-g} \left\{ \frac{m_p^2}{2} R + \Lambda(t) + c(t) g^{00} + F^{(2)}(\delta g^{00}, \delta K_{\mu\nu}, \delta R_{\mu\nu\sigma\lambda}; t) \right\}$$

- Use Stuckelberg trick to restore quintessence field

Matter Couplings - Conformal Coupling

- Metric which the matter couples to can be conformally scaled

$$S_m [e^{\alpha(\phi)} g_{\mu\nu}, \{\psi\}]$$

Extend EFT of Inflation by working in Jordan frame

$$S = \int d^4x \sqrt{-g} \left\{ f(t) \frac{m_p^2}{2} R + \Lambda(t) + c(t) g^{00} + F^{(2)}(\delta g^{00}, \delta K_{\mu\nu}, \delta R_{\mu\nu\sigma\lambda}; t) \right\} + S_m [g_{\mu\nu}, \{\psi\}]$$

Matter Couplings - Stress Energy Tensor

- Stress-Energy Tensor terms need some representation in EFT of Inflation

Extra terms describe any stress-energy tensor dependency

$$S = \int d^4x \sqrt{-g} \left\{ f(t) \frac{m_p^2}{2} R + \Lambda(t) + c(t) g^{00} + g(t) T^{00} + h(t) T \right. \\ \left. + F^{(2)}(\delta g^{00}, \delta K_{\mu\nu}, \delta R_{\mu\nu\sigma\lambda}, \delta T_{\mu\nu}; t) \right\} + S_m[g_{\mu\nu}, \{\psi\}]$$

Summary

- Have constructed a general effective field theory to describe dark energy
- Framework for investigating perturbative behavior is in progress
- Hope to constrain parameters in general descriptions, based on cosmological history and the behavior of cosmic perturbations